

## Weakly relativistic dielectric tensor for arbitrary wavenumbers

Francesco Volpe

*Max-Planck-Institut für Plasmaphysik, EURATOM-Association, Greifswald, Germany*

In approximating Trubnikov's dielectric tensor  $\varepsilon$  [1], the  $\beta_T = \sqrt{2k_B T/mc^2} \ll 1$  limit is often accompanied by the assumptions of quasi-perpendicular incidence [2-11] and small [7-10] or large [11] Larmor radius parameter  $\lambda = \frac{1}{2} \left( \frac{N_\perp \beta_T}{Y} \right)^2$ , where  $N_\perp$  is the refractive index perpendicular to an external magnetic field of normalised magnitude  $Y = \omega_c/\omega$ . On the other hand, electron Bernstein waves (EBWs) are characterised by  $\lambda \gtrsim 1$  and are excited/detected at finite  $N_\parallel$  [12-15], or tend to develop finite  $N_\parallel$  anyway [16]. To describe the propagation and absorption of EBWs and of oblique electromagnetic electron cyclotron waves of arbitrary  $N_\parallel$ , a mildly relativistic formulation of  $\varepsilon$  is derived here with no assumptions on  $N_\parallel$  and  $N_\perp$  except for those which indirectly follow from  $\beta_T \ll 1$ . To start with, the steady-state solution of the linearized Vlasov-Maxwell problem for a uniformly magnetized plasma is [1, 3]:

$$\varepsilon_{ij} = \delta_{ij} + \frac{2i}{\beta_T^2} \frac{\omega_p^2}{\omega^2} \int_0^\infty d\tau T_{jk}^{(1)} \int \frac{d^3\mathbf{u}}{\gamma} u_i u_k f(\mathbf{u}) \exp(i\gamma\tau - i\mathcal{N} \cdot \mathbf{u}) \quad (1)$$

where the convention of implicitly summing on repeated indices is adopted,  $\omega_p$  is the plasma frequency, the time  $\tau$  is renormalized to the wave period  $\omega^{-1}$ ,  $\mathbf{u} = \mathbf{p}/mc$  is the normalized momentum,  $\gamma = \sqrt{1 + \mathbf{u}^2}$  the Lorentz factor and

$$T^{(1)} = \begin{pmatrix} \cos Y\tau & -\sin Y\tau & 0 \\ \sin Y\tau & \cos Y\tau & 0 \\ 0 & 0 & 1 \end{pmatrix}. \quad (2)$$

The axes are chosen to yield  $\mathbf{N} = (N_\perp, 0, N_\parallel)$ . Finally,

$$\mathcal{N}_x = \frac{1}{Y} N_\perp \sin Y\tau, \quad \mathcal{N}_y = \frac{1}{Y} N_\perp (\cos Y\tau - 1), \quad \mathcal{N}_z = N_\parallel \tau \quad (3)$$

and, for a thermal relativistic plasma, the distribution function is the relativistic maxwellian

$$f(\mathbf{u}) = \frac{\exp(-2\gamma/\beta_T^2)}{2\pi\beta_T^2 K_2(2/\beta_T^2)} \quad (4)$$

where  $K_2$  is the modified Bessel function of the second kind of order 2, also known as MacDonald function [18]. Integration over momenta gives [1]:

$$\varepsilon_{ij} = \delta_{ij} + 4i \frac{\omega_p^2}{\omega\omega_c} \frac{1}{\beta_T^4 K_2} \int_0^\infty d\tau \left[ T_{ij}^{(1)} \frac{K_2(R^{1/2})}{R} - T_{ij}^{(2)} \frac{K_3(R^{1/2})}{R^{3/2}} \right] \quad (5)$$

where  $R = 4b^2 + \mathcal{N}^2$ ,  $b = \frac{1}{\beta_T^2} - i\frac{\tau}{2}$  and  $T_{ij}^{(2)} = \mathcal{N}_i T_{jk}^{(1)} \mathcal{N}_k$ .

The limit  $\beta_T \rightarrow 0$  implies  $b \rightarrow \infty$ , that justifies the asymptotic expansions [18]

$$\frac{K_2(R^{1/2})}{R} \simeq e^{-R^{1/2}} \sqrt{\frac{\pi}{2}} \left[ \frac{1}{R^{5/4}} + \frac{15}{8} \frac{1}{R^{7/4}} \right] \quad (6)$$

$$\frac{K_3(R^{1/2})}{R^{3/2}} \simeq e^{-R^{1/2}} \sqrt{\frac{\pi}{2}} \left[ \frac{1}{R^{7/4}} + \frac{35}{8} \frac{1}{R^{9/4}} \right] \quad (7)$$

and their further approximation to the highest orders in  $b$ . This leads to:

$$\varepsilon_{ij} = \delta_{ij} + i \frac{\omega_p^2}{\omega^2} \frac{1}{\beta_T^4 K_2(2\beta_T^{-2})} \frac{\pi^{1/2}}{2} \int_0^\infty d\tau \frac{e^{-2b} e^{-\mathcal{N}^2/4b}}{16b^{11/2}} Q_{ij} \quad (8)$$

where

$$Q_{ij} = T_{ij}^{(1)} b(16b^2 + 15b + 4\mathcal{N}^2) - T_{ij}^{(2)} (8b^2 + 17.5b + \mathcal{N}^2) \quad (9)$$

Note that approximating the exact integral, eq.5, is more rigorous than Taylor-expanding  $\gamma$  up to the second order in  $u$  in the integrand in eq.1, which gives the same as eqs.8-9 but with slightly different coefficients for eq.9 [20].

After the changes of variables  $\beta_T^2 \tau/2 \rightarrow t$  and  $2Y/\beta_T^2 \rightarrow y$  and after expanding:

$$\exp \left[ \frac{\lambda \cos yt}{1 - it} \right] = \sum_{p=0}^{\infty} \sum_{q=0}^{\infty} \left[ \frac{\lambda}{2(1 - it)} \right]^{p+q} \frac{e^{i(p-q)yt}}{p!q!} = \sum_{n=-\infty}^{\infty} \sum_{m=|n|}^{\infty} \frac{(\lambda/2)^m}{\left(\frac{m+n}{2}\right)! \left(\frac{m-n}{2}\right)!} \frac{e^{inyt}}{(1 - it)^m}, \quad (10)$$

with the double sum restricted to even values of  $m+n$ , all integrals in eq.8 take the form

$$\mathcal{Q}_{mn,ij} = e^{-\lambda} \beta_T^6 \int_0^\infty dt \frac{Q_{ij}}{(1 - it)^{m+11/2}} \exp \left[ \frac{2}{\beta_T^2} (nY + 1) it - \frac{\lambda + N_{\parallel}^2 t^2 / \beta_T^2}{1 - it} \right] \quad (11)$$

involving generalized Shkarofsky functions [9, 21, 22]

$$\mathcal{F}_{q,r}(z, a) = -i \int_0^\infty \frac{(it)^r}{(1 - it)^q} \exp \left[ izt - \frac{at^2}{1 - it} \right] dt, \quad (12)$$

of shifted arguments, that will be omitted for brevity and summarized by an index for the harmonic number:

$$\mathcal{F}_{q,r,n} = \mathcal{F}_{q,r} \left( \frac{2}{\beta_T^2} (nY + 1) - \lambda, \frac{N_{\parallel}^2}{\beta_T^2} - \lambda \right). \quad (13)$$

Substituting in eq.11 an approximation up to  $\mathcal{O}(\beta_T^4)$  of eq.9 and replacing, for consistency,  $K_2$  at denominator of eq.1 with its asymptotic limit, yield the following expression for the slightly relativistic dielectric tensor:

$$\varepsilon_{ij} = \delta_{ij} + i \frac{\omega_p^2}{\omega^2} \frac{e^\lambda}{16 + 15\beta_T^2} \sum_{n=-\infty}^{\infty} \sum_{m=|n|}^{\infty} \frac{(\lambda/2)^m \mathcal{Q}_{mn,ij}}{\left(\frac{m+n}{2}\right)! \left(\frac{m-n}{2}\right)!} \quad (14)$$

where, as usual, the summation is restricted to even values of  $m + n$ , and

$$\mathcal{Q}_{mn,11} = 8i(1 - N_{\parallel}^2 O_{2,2,0}) \mathcal{F}_{m+\frac{5}{2},0,n\pm 1}^+ + \frac{15}{2} i \beta_T^2 \mathcal{F}_{m+\frac{7}{2},0,n\pm 1}^+ \quad (15)$$

$$+ i \beta_T^2 \frac{N_{\perp}^2}{Y^2} (2 - N_{\parallel}^2 O_{2,2,0}) \left( \mathcal{F}_{m+\frac{7}{2},0,n\pm 2}^+ - 2 \mathcal{F}_{m+\frac{7}{2},0,n}^+ \right) \quad (16)$$

$$\mathcal{Q}_{mn,22} = 8i(1 - N_{\parallel}^2 O_{2,2,0}) \mathcal{F}_{m+\frac{5}{2},0,n\pm 1}^+ + \frac{15}{2} i \beta_T^2 \mathcal{F}_{m+\frac{7}{2},0,n\pm 1}^+ \quad (17)$$

$$- i \beta_T^2 \frac{N_{\perp}^2}{Y^2} (2 - N_{\parallel}^2 O_{2,2,0}) \left( \mathcal{F}_{m+\frac{7}{2},0,n\pm 2}^+ - 4 \mathcal{F}_{m+\frac{7}{2},0,n\pm 1}^+ + 6 \mathcal{F}_{m+\frac{7}{2},0,n}^+ \right) \quad (18)$$

$$\mathcal{Q}_{mn,21} = 8(1 - N_{\parallel}^2 O_{2,2,0}) \mathcal{F}_{m+\frac{5}{2},0,n\pm 1}^- + \frac{15}{2} \beta_T^2 \mathcal{F}_{m+\frac{7}{2},0,n\pm 1}^- \quad (19)$$

$$+ \beta_T^2 \frac{N_{\perp}^2}{Y^2} (2 - N_{\parallel}^2 O_{2,2,0}) \left( \mathcal{F}_{m+\frac{7}{2},0,n\pm 2}^- - 2 \mathcal{F}_{m+\frac{7}{2},0,n\pm 1}^- \right) \quad (20)$$

$$\mathcal{Q}_{mn,33} = \frac{16i}{\beta_T^2} N_{\parallel}^2 (2 - N_{\parallel}^2 O_{2,2,0}) \mathcal{F}_{m+\frac{7}{2},2,n} + 2i(8 + 27N_{\parallel}^2 O_{2,2,0}) \mathcal{F}_{m+\frac{5}{2},0,n} \quad (21)$$

$$+ \frac{15i \beta_T^2 \mathcal{F}_{m+\frac{7}{2},0,n} - 4i \beta_T^2 \frac{N_{\perp}^2}{Y^2} N_{\parallel}^2 \left( \mathcal{F}_{m+\frac{11}{2},2,n\pm 1}^+ - 2 \mathcal{F}_{m+\frac{11}{2},2,n}^+ \right)}{\beta_T^2} \quad (22)$$

$$\mathcal{Q}_{mn,13} = 4i(2 - N_{\parallel}^2 O_{2,2,0}) \frac{N_{\perp}}{Y} N_{\parallel} \mathcal{F}_{m+\frac{7}{2},1,n\pm 1}^- + \frac{35}{2} i \beta_T^2 \frac{N_{\perp}}{Y} N_{\parallel} \mathcal{F}_{m+\frac{9}{2},1,n\pm 1}^- \quad (23)$$

$$\mathcal{Q}_{mn,23} = 8(2 - N_{\parallel}^2 O_{2,2,0}) \frac{N_{\perp}}{Y} N_{\parallel} \mathcal{F}_{m+\frac{7}{2},1,n} + 35 \beta_T^2 \frac{N_{\perp}}{Y} N_{\parallel} - 2 \mathcal{F}_{m+\frac{9}{2},1,n} + i \mathcal{Q}_{mn,13} \quad (24)$$

$\mathcal{Q}_{mn,12} = -\mathcal{Q}_{mn,21}$ ,  $\mathcal{Q}_{mn,31} = \mathcal{Q}_{mn,13}$  and  $\mathcal{Q}_{mn,32} = -\mathcal{Q}_{mn,23}$ . The compact notations

$$\mathcal{F}_{q,r,n\pm p}^+ = \mathcal{F}_{q,r,n+p} + \mathcal{F}_{q,r,n-p} \quad (25)$$

$$\mathcal{F}_{q,r,n\pm p}^- = \mathcal{F}_{q,r,n+p} - \mathcal{F}_{q,r,n-p} \quad (26)$$

were employed for some frequently occurring sums and differences and an operator on the subscripts  $q$  and  $r$ , defined by  $O_{2,2,0} \mathcal{F}_{q,r,n} = \mathcal{F}_{q+2,r+2,n}$ , was introduced.

In eq.13 dispersion was found to depend on Larmor radius, in agreement with relativistic eq.1. The shift of generalized Shkarofsky functions' arguments  $z$  and  $a$  by an amount  $\lambda$  can be interpreted as a linear finite-Larmor-radius correction to the resonance condition through  $z = 2(nY + 1)/\beta_T^2$  and to its width in inhomogeneous plasmas, through the ratio of Doppler to relativistic width,  $a = N_{\parallel}/\beta_T$ . The physical meaning of  $z \rightarrow z - \lambda$  is the well-known relativistic downshift.

At this point it should be remembered that it is customary to treat as a constant the Lorentz factor  $\gamma$  in the denominator in the integrand of eq.1 and to Taylor-expand only the factor  $\gamma$  in the exponent, on the ground that relativistic corrections in the denominator have a comparatively small effect on the integral. This is equivalent to approximating  $Q_{ij} \simeq (T_{ij}^{(1)} 2b - T_{ij}^{(2)}) 8b^2$

i.e. to suppressing all underlined terms in eqs.15-24, which is only partly legitimate under the assumption  $N_{\parallel} \ll \beta_T$  that, by the way, was not adopted in the present work.

The double sum in eq.14 might look computationally expensive, but functions  $\mathcal{F}$  of different  $q$  (thus,  $\mathcal{Q}$  of different  $m$ ) can recursively be related to one another [2, 5, 8, 10, 21]. Note also that simple Shkarofsky functions of  $r = 0$  are the most widely used in eqs.15-24. Generalized functions  $\mathcal{F}_{q,r}$  with  $r \neq 0$  can be expressed in terms of  $\mathcal{F}_q$  [17] and this be related to the classical dispersion function  $Z$  [8]. The latter coincides with the Dawson's integral, for the evaluation of which plenty of routines are available. Besides, for relativistically broadened but well-resolved lines, the double sum can be restricted to diagonal terms ( $m = n$ ).

An alternative form of eq.14 can be obtained by recognizing, in the sum over  $m$  in eq.10, the generating function for the modified Bessel function of the first kind,  $I_n$ . This generalizes a standard Bessel function identity utilized in the derivation of the warm non-relativistic dielectric tensor [19]. However, it does so by replacing  $\lambda \rightarrow \lambda/(1-it)$ , i.e. by introducing a time dependence for  $I_n$  that cannot be factored out of the integral anymore. The resulting expression,

$$\varepsilon_{ij} = \delta_{ij} + 2i \frac{\omega_p^2}{\omega^2} \frac{\beta_T^4}{16 + 15\beta_T^2} \sum_{n=-\infty}^{\infty} \int_0^{\infty} dt \frac{Q_{ij}}{(1-it)^{11/2}} \exp \left[ \frac{2}{\beta_T^2} (nY + 1) - \frac{\lambda + N_{\parallel}^2 t^2 / \beta_T^2}{1-it} \right] I_n \left( \frac{\lambda}{1-it} \right) \quad (27)$$

involves integrals more complicated than eq.11. On the other hand, it avoids the nuisance of the double sum and can be useful when a high  $m$  is required for convergence.

## References

- [1] B.A. Trubnikov, Plasma Physics and the Problem of Controlled Thermonuclear Reactions (Pergamon Press, New York, 1959), vol.III, p.122
- [2] I. P. Shkarofsky, J. Plasma Phys.**35**, 319 (1986)
- [3] P. H. Yoon and R. C. Davidson, J. Plasma Phys.**43**, 269 (1990)
- [4] P. A. Robinson, Phys. Fluids **31**, 107 (1988)
- [5] D. G. Swanson, Plasma Phys. Control. Fusion **44**, 1329 (2002)
- [6] A. D. Piliya *et al.*, Plasma Phys. Control. Fusion **45**, 1309 (2003)
- [7] I. P. Shkarofsky, Phys. Fluids **9**, 561 (1966)
- [8] V. Krivenski and A. Orefice, J. Plasma Phys.**30**, 125 (1983)
- [9] P. A. Robinson, J. Math. Phys.**27**, 1206 (1986)
- [10] P. A. Robinson, J. Math. Phys.**28**, 1203 (1987)
- [11] E. Lazzaro and A. Orefice, Phys. Fluids **23**, 2330 (1980)
- [12] H. P. Laqua *et al.*, Phys. Rev. Lett.**78**, 3467 (1997)
- [13] H. P. Laqua *et al.*, Phys. Rev. Lett.**81**, 2060 (1998)
- [14] F.Volpe *et al.* Rev. Sci. Instrum.**74**, 1409 (2003)
- [15] H. P. Laqua *et al.*, Phys. Rev. Lett. **90**, 75003, (2003)
- [16] C. B. Forest *et al.*, Phys. Plasmas **7**, 1352 (2001)
- [17] A. C. Airoidi and A. Orefice, J. Plasma Phys.**27**, 515 (1982)
- [18] M. Abramowitz and I. A. Stegun, *Handbook of Mathematical Functions*, 1974
- [19] M. Brambilla, *Kinetic Theory of Plasma Waves* (Oxford Univ.Press, 1998)
- [20] F. Volpe and N. M. Marushchenko, Proc.EC-14 Workshop, 9-12 May 2006 Santorini (Greece)
- [21] N. M. Temme *et al.*, Astrophys. and Space Sc.**194**, 173 (1992)
- [22] D. B. Melrose, *Quantum Plasmadynamics*, in preparation
- [23] P. A. Robinson, J. Plasma Phys.**37**, 435 (1987)
- [24] P. A. Robinson, J. Plasma Phys.**37**, 449 (1987)