

Onset of Driven Collisionless Reconnection in Magnetized Pair Plasmas

C. Granier¹, D. Grošelj¹, L. Comisso^{2,3}, and F. Bacchini^{1,4}

¹Centre for mathematical Plasma Astrophysics, Department of Mathematics, KU Leuven, Belgium

²Department of Physics, Columbia University, New York, NY, USA

³Department of Astronomy, Columbia University, New York, NY, USA

⁴Royal Belgian Institute for Space Aeronomy, Belgium

Abstract

We present 2D particle-in-cell simulations of current-sheet formation, disruption, and nonlinear merging in a strongly magnetized electron–positron plasma. By driving two force-free flux tubes of radius R together at speed v_{push} , we observe an exponential collapse. Reconnection onset occurs when the half-thickness reaches $a \simeq d_e$, after

$$t_{\text{onset}} \simeq \frac{R}{v_A + v_{\text{push}}} \left[1 + \ln \frac{R v_A / (v_A + v_{\text{push}})}{d_e} \right].$$

Plasmoid chains emerge only for sheet aspect ratios $A = l/a \gtrsim 30$ (where l is the sheet length and a its thickness). In the subsequent merging phase lasting $\Delta t_{\text{merg}} \sim 2R/c$, roughly 30% of the reconnecting magnetic energy is dissipated at a normalized rate close to ~ 0.1 , largely independent of drive speed or magnetization. These quantitative scalings provide benchmarks for flare models in compact-object magnetospheres.

Introduction Magnetic reconnection, which explains how solar flares release energy [1], is also used to account for bright flares in pulsar wind nebulae [2] and sudden bursts around neutron stars and black holes [3]. In these extreme environments, reconnection happens so fast that it reaches nearly the speed of light [4]. Traditional kinetic studies often begin from preformed Harris sheets or coalescing islands, thereby missing the self-consistent formation of the reconnecting layer. Here, we drive two force-free flux tubes together in a controlled 2D geometry to capture the full lifecycle of the sheet: from formation and kinetic tearing to plasmoid onset and final tubular merger.

Methodology & Evolution Phases We employ the Tristan-MP v2 PIC code in 2D, initializing two Lundquist-type force-free flux tubes [5, 6] of radius R in a cold pair plasma ($\sigma_0 = B_0^2/4\pi n_0 m_e c^2$ from 25 to 800). The tubes are separated by $2R$ in a $4R \times 4R$ domain resolved at $\Delta x = d_e/5$, with $R/d_e \in [50, 1600]$. A uniform inflow v_{push} (so $t_{\text{drive}} = R/v_{\text{push}}$) is imposed via $\mathbf{E} = -\mathbf{v}_{\text{push}} \times \mathbf{B}/c$. We identify three principal phases:

- *Phase I – Sheet Formation:* The current sheet forms, while small-amplitude tearing modes ($k_x d_e \lesssim 1$) grow linearly.

- **Phase II – Plasmoid Formation:** The sheet collapses and nonlinear tearing leads to a chain of plasmoids when $A = \ell/a \gtrsim 30$. The normalized reconnection rates are around 0.002 and 0.02.
- **Phase III – Merging:** Flux tubes coalesce rapidly over $\Delta t_{\text{merg}} \sim 2R/c$, with a normalized reconnection rate close to 0.1.

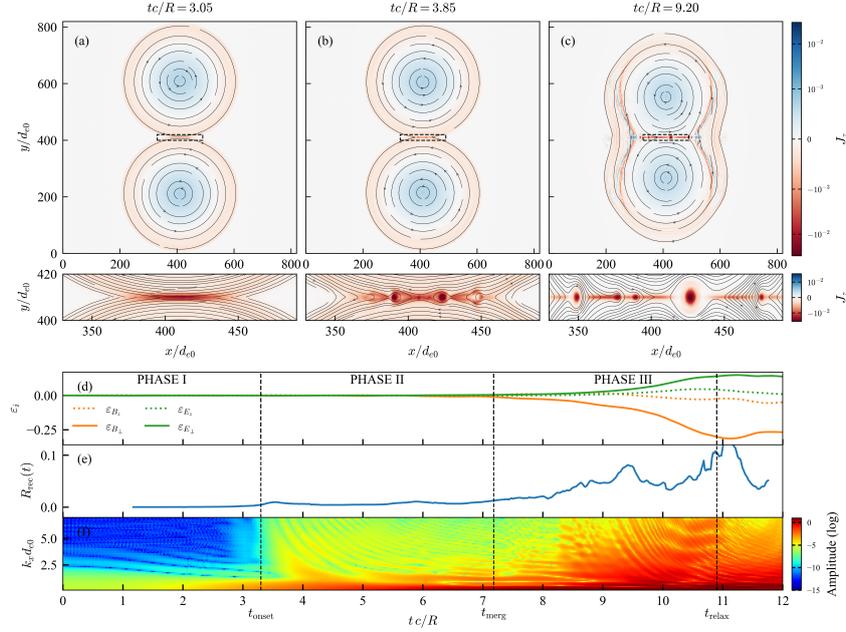


Figure 1: (a)–(c): Contours of out-of-plane current density J_z (color scale) and magnetic field lines (black) during the evolution of the flux tubes. The dashed black boxes mark the regions shown in the zoomed-in view below. (d): Temporal evolution of the energy components, $\epsilon_i = (E_i(t) - E_i(0))/E_{B_\perp}(0)$. (e): Time evolution of the normalized reconnection rate. (f): k_x -spectra of the magnetic flux function averaged inside the current sheet (over $10d_{e0}$ along y). The color-coded intensity represents the logarithmic amplitude of the wavenumbers. Key time thresholds (t_{onset} , t_{merg} and t_{relax}) are indicated by vertical dashed lines. This simulation was run with $v_{\text{push}} = 0.05c$ ($t_{\text{drive}} = 20R/c$), $\sigma_0 = 50$ ($\sigma_{\text{in}} = 8$), and $R/d_{e0} = 205$.

Linear Tearing-Mode Growth In our simulations, during Phase I the modes satisfying $k_x d_e \lesssim 1$ exhibit a linear growth rate. Figure 2 plots the logarithmic amplitude of several Fourier modes versus time, confirming that the linear stage ends precisely when the sheet thickness approaches the critical scale a_K .

Onset of Reconnection We derive the onset time by matching the linear thinning and exponential collapse regimes. For $t < t_K$, under the converging inflow v_{push} , the two Lundquist tubes of radius R develop a sheet whose half-width $a(t)$ follows first a kinematic thinning (frozen-in drive), $a(t) = R - v_{\text{push}}t$ with t_K defined as the moment when pinch acceleration $\sim (v_A^2)/a$ becomes comparable to the imposed inflow. Beyond $t > t_K$, the self-pinch ($\mathbf{J} \times \mathbf{B}$ Lorentz force) drives the thinning and yields $da/dt = -a/\tau$ with $\tau = a_K/v_A$, so $a(t - t_K) = a_K e^{-(t-t_K)/\tau}$. We equate the external driving time $t_K = (R - a_K)/v_{\text{push}}$ with the Alfvén time $t_A = a_K/v_A$. This

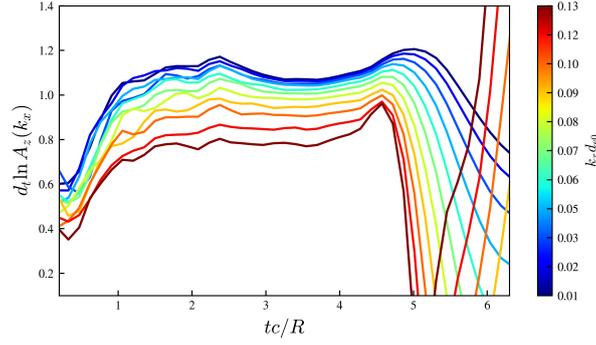


Figure 2: Evolution of tearing-mode amplitudes (log scale) for modes $k_x d_e = 0.5, 0.8, 1.0$ during Phase I. The transition to nonlinear growth coincides with the onset time t_{onset} .

yields $t_K \simeq R/v_A + v_{\text{push}}$, and $a_K \simeq R v_A / v_A + v_{\text{push}}$. Setting $a(t_{\text{onset}}) \simeq d_e$ gives

$$t_{\text{onset}} \simeq \frac{R}{v_A + v_{\text{push}}} \left[1 + \ln \frac{a_K}{d_e} \right]. \quad (1)$$

Figure 3 shows excellent agreement between this prediction and simulation results across a broad range of R/d_e , v_{push} , and σ_{in} . The weak dependence on σ_{in} arises because $v_A \simeq c$ when $\sigma_{\text{in}} \gg 1$.

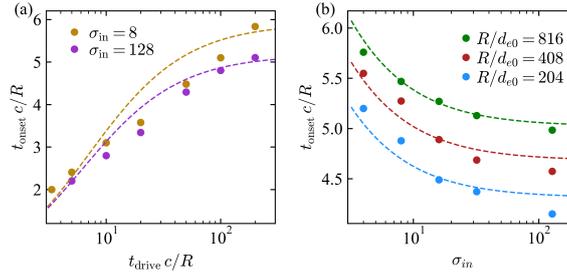


Figure 3: Normalized onset time $t_{\text{onset}}c/R$ vs. driving timescale $t_{\text{drive}}c/R$ for two magnetizations ($\sigma_{\text{in}} = 8, 128$). Symbols: simulation; lines: theoretical prediction.

Reconnection Rate & Dissipation We measure the instantaneous reconnection rate via

$$R_{\text{rec}}(t) = \frac{1}{v_A B_{\text{xup}}} \frac{d}{dt} [\max(A_z) - \min(A_z)], \quad (2)$$

where A_z is the out-of-plane component of magnetic vector potential in the midplane between the two tubes (i.e., along the current sheet), and B_{xup} and v_A are the upstream reconnecting field and Alfvén speed (based on B_{xup}). Equation (2) measures how fast magnetic flux is transferred across the separatrix.

During rapid merging, the rate saturates near $R_{\text{rec}} \sim 0.075$, and the energy-based rate

$$\mathcal{R}_{\text{diss}} = \frac{R}{cE_{B\perp}(0)} \frac{d[-E_{B\perp}(t)]}{dt} \approx 2R_{\text{rec}}, \quad (3)$$

which is the normalized rate of in-plane magnetic energy loss (the factor 2 comes from Poynting inflows on both sides feeding the X-point).

Figure 4 shows $R_{\text{rec}} \rightarrow 0.1$ in Phase III, independent of drive. The merging duration $\Delta t_{\text{merg}} \approx 2R/c$ yields $\langle \mathcal{R}_{\text{diss}} \rangle \approx 0.19$.

Conclusions Our 2D PIC simulations of driven collisionless reconnection in strongly magnetized pair show that the onset time follows a simple analytic scaling $\propto R/(c + v_{\text{push}})[1 + \ln(Rc/d_e)]$, plasmoids require sheet aspect ratios $\gtrsim 30$, and the merging phase dissipates $\sim 30\%$ of reconnecting energy at a universal rate close to ~ 0.1 (see arXiv [6] for full details).

References

- [1] Parker, E. N. 1957, *J. Geophys. Res.*, 62, 509.
- [2] Uzdensky, D. A., Cerutti, B., & Begelman, M. C. 2011, *ApJL*, 737, L40.
- [3] Thompson, C. 1994, *MNRAS*, 270, 480.
- [4] Sironi, L., Uzdensky, D. A., & Giannios, D. 2025, *ARA&A*, 63, doi:10.1146/annurev-astro-020325-115713.
- [5] Lyutikov, M., Sironi, L., Komissarov, S. S., & Porth, O. 2017, *J. Plasma Phys.*, 83, 635830602, doi:10.1017/S002237781700071X.
- [6] Granier, C., Grošelj, D., Comisso, L., & Bacchini, F. 2025, *ApJ*, submitted; arXiv:2506.06059.

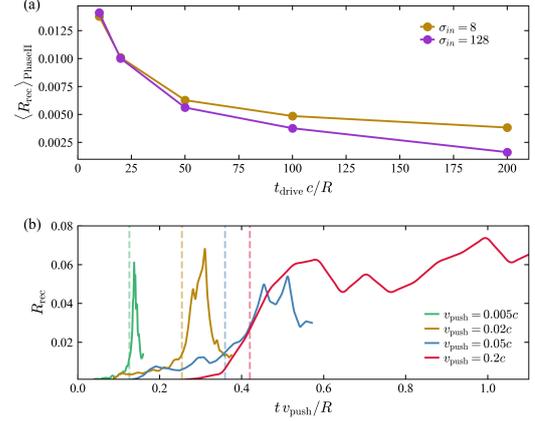


Figure 4: Reconnection rate $R_{\text{rec}}(t)$ for various push velocities v_{push}/c . The dashed line marks the beginning of Phase III.

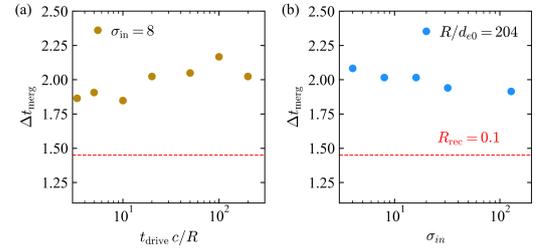


Figure 5: (a) Normalized merging duration $\Delta t_{\text{merg}} c/R$ versus driving timescale for $\sigma_{\text{in}} = 8$ and $R/d_{e0} = 204$. (b) $\Delta t_{\text{merg}} c/R$ versus in-plane magnetization for $t_{\text{drive}} c/R = 50$. Start/end defined at 1%/30% energy dissipated.